

Review

Younger gabbros from south Sinai: Petrology, geochemistry and petrogenetic aspects

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Five minor isolated intrusions of younger gabbros crop out in Southeastern-, Central- and Southeastern Sinai are presented. The gabbros range in composition from normal, pyroxene-hornblende, hornblende, and leucogabbro and pyroxenite (low Ti content) to normal, pyroxene-hornblende, hornblende, olivine and uralitized gabbro, anorthosite and hornblendite (high Ti content). They are derived from two magma types (tholeiitic and calc-alkaline) closely related in time and space, but with different geochemical trends. The tholeiitic younger gabbros (THYG) at W. Tweiba, W. Rahaba and El-Khamila areas are generally characterized by high FeO^t, TiO₂, MnO, Zr, Y, Nb, Th and total REE contents as compared with the calc-alkaline younger gabbros (CAYG) at Imliq and W. Nakhil areas. The parental magma of the THYG was controlled by multistage processes including varying degrees of crystal fractionation and differentiation with prominent crustal contamination. The parental magma of the CAYG was controlled by differentiation and crystallization of anhydrous minerals which accompanied by high oxygen fugacity and water activity. The present gabbros were suffered a minor crustal materials involvement. The gabbros of all localities have LREE enrichment; HREE depletion and a small positive Eu anomaly suggest their moderate differentiation. The REE patterns reveal the dominant role of plagioclase over the pyroxene. The absence of a positive Eu anomaly of the cumulus plagioclase is due to pressure fractionation before emplacement. A slight positive Eu anomaly in some gabbros is attributed to plagioclase accumulation. The diversities of both kinds of gabbros are ascribed to the varying degrees of partial melting, crustal contamination and/or subduction related contamination by subcontinental lithosphere. The existence of the two magma types reveals a heterogeneous mantle or upper mantle-lower crust source. They were emplaced in continental crust and tend to be formed during the final stage of arc to active continental margin.

Keywords: Younger gabbros, geochemistry, petrogenesis, five occurrences, south Sinai, Egypt.

INTRODUCTION

The Egyptian basement complex splits into four main groups (El-Gaby, 2005): 1) Pre-Pan-African rocks comprising deformed granites, migmatites, gneisses and high-grade Barrovian-type metamorphites, 2) Pan-African ophiolite and island arc assemblages thrust onto the old continent causing deformation and cataclasis in the over-ridden infrastructural rocks, 3) Pan-African Cordillera stage comprising calc-alkaline gabbro-diorite complexes, Dokhan Volcanics, Hammamat sediments, calc-alkaline granite series, together with olivine gabbro and related rocks, and 4) post orogenic to anorogenic alkaline to per-alkaline silicic magmatism including alkali feldspar granites, syenites and alkali rhyolites.

The Egyptian gabbros have been recently classified

into three major types (Khalil, 2005): 1) Ophiolitic metagabbros, 2) Arc-related metagabbros and 3) Cordilleran stage gabbros (younger gabbros). The ophiolitic gabbros were termed epidiorites complex (El-Ramly and Akaad, 1960), or metagabbro-diorite complex (Akaad and Essawy, 1964; El-Ramly, 1972) or older metagabbros (Takla et al., 1981). They are regionally metamorphosed and dated 800-730 Ma (Kroner et al., 1990). The island arc metagabbros are of limited distribution and difficult to separate them from the ophiolitic metagabbros. Cordilleran stage gabbros (younger gabbros) are mafic-ultramafic cumulates or intrusive unmetamorphosed rocks (El-Ramly, 1972). They were named fresh, layered younger gabbros (Takla

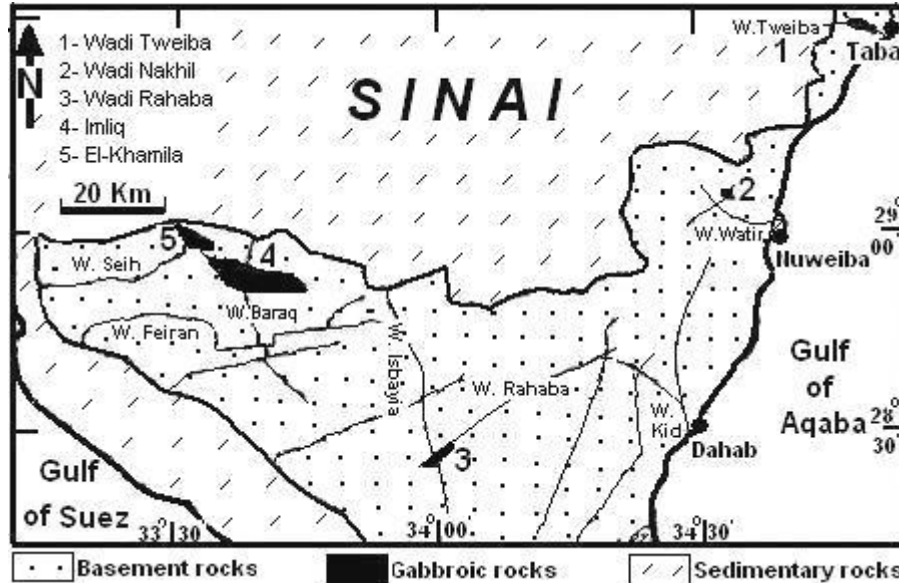


Figure 1. Location map showing the distribution of the studied younger gabbros in south Sinai.

et al., 1981), and considered to be post-tectonic intrusions (655-570 Ma), equivalent to the Andean type calc-alkaline sequence (El-Gaby et al., 1988) or olivine gabbro and related rocks (El-Gaby, 2005).

Gabbroic rocks in Sinai crop out as numerous small masses. El-Metwally (1997) suggested that the Sinai mafic-ultramafic rocks (metagabbro-diorite complex) were intruded in an active continental margin above subduction zone. They were metamorphosed into the greenschist, and locally in the lower amphibolite facies. Basta (1998) concluded that the Sinai gabbros (including Wateir) were crystallized from a hydrous calc-alkaline magma, possibly formed in an active continental margin environment. Kroner et al. (1990) determined the ages of metagabbro-metabasites complex and metadiorite from Taba (nearby Tweiiba) and Aqaba using $^{207}\text{Pb}/^{209}\text{Pb}$ method and obtained ages ranging from 640 to 641 ± 6 Ma. Abdel-Karim (1995) concluded that the K-Ar age dating of the hornblendes and biotites separated from metagabbro-diorite complex from Wadi Baba, SW Sinai gave 794 ± 30 to 667 ± 25 Ma age. Takla et al. (2001) considered the ophiolitic ultramafics of Kabr El-Bonya of SE Sinai (Abdel-Khalek et al., 1997), layered ophiolitic metagabbro-diorite of El-Melhega (Shimron, 1981; Abdel-Khalek et al., 1997) and the metagabbro-diorite of Feiran (Higazy and El-Gammal, 1990), Nisreen (El Tokhi and Katta, 1993) and Saal areas (El-Metwally, 1997) all together as intrusive younger gabbros. Tweiiba gabbros were previously described as metagabbro-diorite complex of tholeiitic affinity, intruded in a continental within-plate tectonic setting and metamorphosed into the middle amphibolite facies (Abu El-Enen, 1995). Takla et al. (1991) concluded that the hornblende gabbro of Wadi Watir (including Wadi Nakhil) was affiliated to the

younger gabbros. The gabbroic rocks of Rahaba were previously described as mafic-ultramafics, composed of pyroxene metagabbro in the inner part of the pluton, and hornblende in the periphery or the sheared mylonitic zones (Soliman, 1997). El-Metwally (1986) studied the rocks crop out in the El-Sammraa (Imliq) and Wadi Seih (El-Khamila) and considered them as two mafic-ultramafic intrusions. The former intrusion composed of layered wehrlite, clinopyroxenite and gabbros, while the second one comprised cumulate gabbros, pyroxenites and anorthosites, both of them are considered as a non-ophiolitic assemblage. On the other hand, Mehanna et al., (2004) concluded that the mafic-ultramafic rocks of Imliq area represented by serpentinized peridotites and metagabbros belong to an ophiolite assemblage. Petrographic and chemical characterization of Fe-Ti oxides and sulfides hosted in the present mafic intrusions were given by Abdel-Karim (2009).

General geology and petrography

The studied gabbroic rocks of Sinai occur in three main provinces: Southeastern Sinai (Tweiiba and Nakhil), Central Sinai (Rahaba) and Southeastern Sinai (Imliq and El-Khamila) (Figure 1).

The Tweiiba gabbros form a small elongated mass, which occupies an area of about 3.2 km^2 . It exhibits a sharp intrusive contact against the adjacent granitoid rocks (Figure 2a) and schist. The rock is medium grained with massive and equi-dimensional textures. It comprises pyroxene hornblende gabbros in the inner part of the intrusion, and hornblende gabbros in the periphery. Petrographically, the hornblende gabbros are made up of

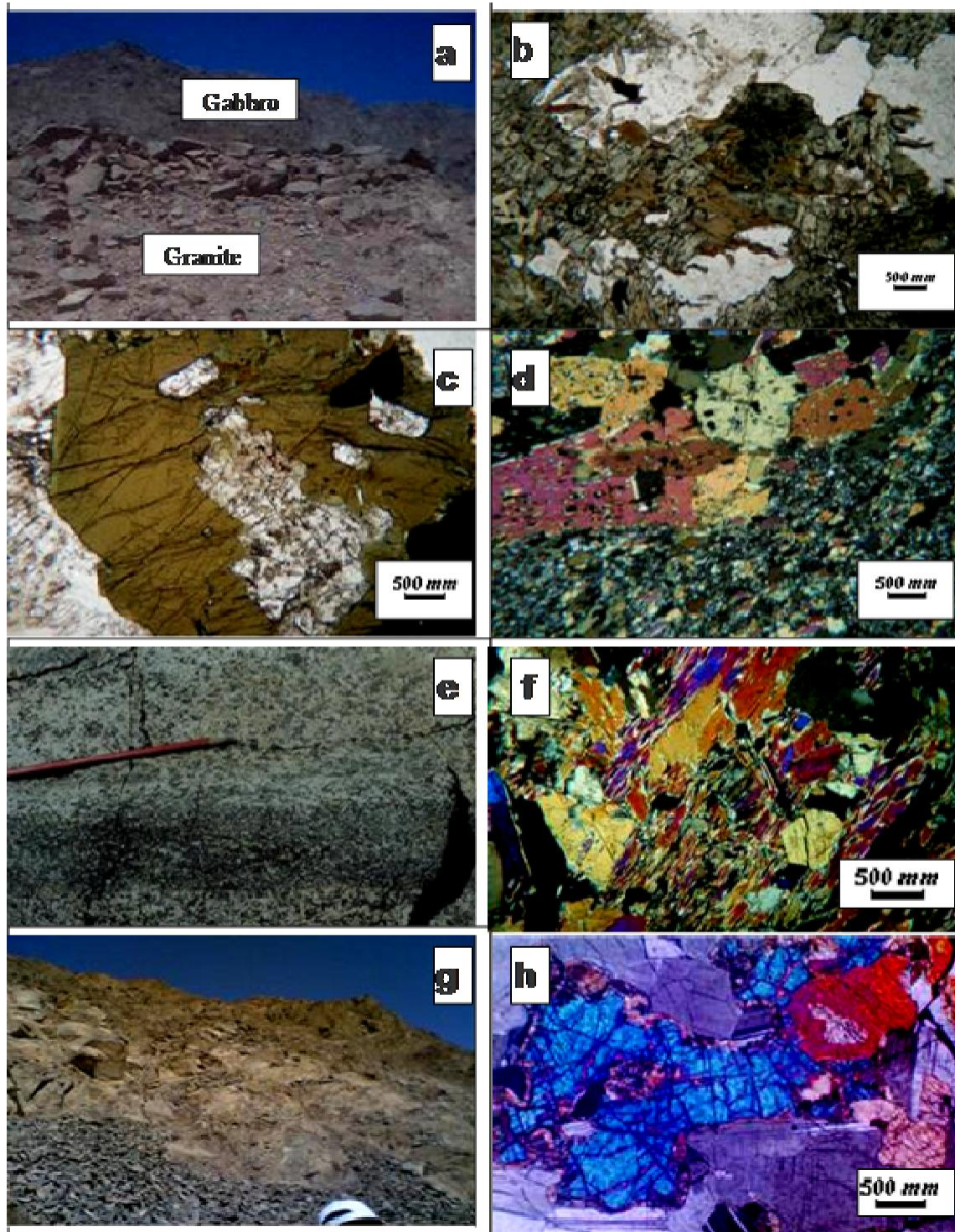


Figure 2. Field photos and photomicrographs showing: a) sharp contact between gabbro and granite, W . Toweiba, b) partial alteration of oxyhornblende to biotite and opaques, Hornblende gabbro, W . Tweiba, c) inclusion of plagioclase and opaques in pyroxene, Normal gabbro, W . Nakhil, d) abundance of cumulus hornblende megacrysts , Hornblendite, W . Rahaba, e) rhythmic layering of gabbro, Imliq, f) dominant cumulus diopside, Clinopyroxenite, Imliq, g) weathering surface of gabbro, El-Khamila; h) olivine is replaced by pyroxene or mantled by amphibole, Olivine gabbro, El-Khamila.

cumulus hornblende and plagioclase, with subordinate amounts of biotite, quartz with accessory Fe oxides, sulfides and apatite. The hornblende is fresh, partly

altered to biotite and iron oxides (Figure 2b) or recrystallized to fine grained aggregates of the same minerals. The pyroxene hornblende gabbros are

composed essentially of cumulus hornblende, pyroxene and plagioclase with subordinate amounts of quartz, biotite, and accessory magnetite, ilmenite and apatite.

Wadi Nakhil gabbros form an oval-shape mass (with an area of about 1.0 km²). The mass occurs as large xenoliths or as roof pendants within the neighboring Younger Granites. The rock is fresh, massive and coarse to rarely medium grained and comprises pyroxene-hornblende gabbros and normal gabbros. Petrographically, the hornblende gabbros are essentially composed of cumulus plagioclase (An₅₀₋₅₆) and hornblende with subordinate amounts of intercumulus augite. Quartz, biotite, chlorite, apatite and opaques are accessories. The normal gabbros are essentially composed of cumulus plagioclase (An₅₂₋₆₀), augite and intercumulus hornblende with subordinate amounts of biotite. Accessories are quartz, chlorite, apatite and opaques. Augite is partly or totally enclosed variable-size plagioclase forming ophitic and subophitic textures (Figure 2c). The pyroxene-hornblende gabbros are essentially composed of cumulus plagioclase (An₅₀₋₅₅) and pyroxene with subordinate amounts of hornblende. Accessories are apatite, magnetite and titanomagnetite.

Wadi Rahaba gabbros occur as a small elongate mass (about 4.2 km²). It comprises uralitized gabbros and hornblendites. The hornblendites form dyke-like bodies cutting across the uralitized gabbros with sharp contacts and extend for about 15 m long and 2.5 widths. Appinitic gabbro variety is also recorded. The Rahaba gabbros display a sharp intrusive contact against and are located within the adjacent tonalite-granodiorite association. The rock is coarse grained with massive and less common schistose textures. Petrographically, the uralitized gabbros are essentially composed of cumulus augite and hypersthene and plagioclase (An₅₄₋₅₉) with subordinate amounts of pigeonite, quartz and accessory apatite, opaques and chlorite. The hornblendites are mainly composed of cumulus hornblende and plagioclase (An₅₂₋₅₇) with minor pyroxene, biotite, quartz and accessories apatite and iron oxides, as well as chlorite as a retrograde mineral. Hornblende forms sheared megacrysts, usually embedded in a schistose groundmass of hornblende, plagioclase, opaques and titanite (Figure 2d).

Imliq gabbros form an irregular elongated mass (~ 50 km²). The mass intrudes calc-alkaline older granitoids and is intruded by alkaline younger granites. Imliq gabbros mainly comprise pyroxene gabbro and leucogabbro. Lenses and dyke-like bodies of clinopyroxenite (10-25 m long and 1.0-2.5 m width) are observed. It is observed that anorthosite is intruding the gabbro mass. Imliq gabbros are intersected by dyke-swarms of varying composition and styles. Rhythmic layers (a few cm thick) are recorded (Figure 2e). The rock is fresh, medium to very coarse-grained. Petrographically, the normal gabbros are basically

composed of cumulus plagioclase, pyroxene and hornblende. The accessory phase includes quartz, chlorite, titanomagnetite, apatite and ilmenite. Leucogabbros are essentially composed of cumulus augite, hornblende and intercumulus plagioclase with accessory quartz, chlorite, magnetite, ilmenite and apatite. Ophitic and subophitic intergrowths of plagioclase and hornblende are common. Clinopyroxenites are fundamentally composed of cumulus diopside (Figure. 2f) and its alteration products together with minor postcumulus plagioclase (An₈₅₋₈₈). The accessories are magnetite, ilmenite, titanite and chalcopyrite.

El-Khamila gabbros occur as an irregular mass (~ 13 km²). The mass intrudes calc-alkaline older granitoids with a sharp intrusive contact. It sometimes exhibits boulders weathering (Figure 2g). It mainly comprises olivine pyroxene, pyroxene gabbros and less common anorthosite. In places, El-Khamila gabbros exhibit small rhythmic layers (a few cm thick). Igneous lamination is associated with layering. The rock is hard, medium to very coarsegrained, grey to dark green in color. Petrographically, the olivine gabbros are essentially composed of cumulus bytownite (An₆₅₋₇₃), augite and less common olivine. Olivine is replaced by pyroxene or mantled by amphibole (Figure 2h). Hypersthene, magnetite, and ilmenite are accessories. Chlorite is the main secondary mineral. The normal gabbros are made up of cumulus labradorite (An₅₂₋₅₈) and augite with minor intercumulus hornblende and pigeonite. Magnetite, apatite, and titanite are the main accessories. The anorthosites consist of cumulus plagioclase (An₇₅₋₈₅), and minor uralitized diopside. Magnetite and ilmenite are accessories, whereas chlorite and epidote are secondary minerals.

Geochemistry

The geochemistry of the studied younger gabbros is based on twenty-four representative samples. The trace elements were analyzed by using XRF spectrometer. The content of eleven REEs was estimated by ICP Analysis. The analysis has been carried out in the Institute of Mineralogy and crystallography, Vienna University. The analytical results are presented in Table 1.

On the basis of the geochemistry, the Younger Gabbros of Egypt reveal three main concepts:

i) Intrusive rocks of calc-alkaline character and belong to an island arc setting (Takla et al., 1981, Salem et al., 1992, El-Mansey, 1996)

ii) Calc-alkaline to tholeiitic affinity and emplaced into island arc (Sadek, 1994, Moghazi, 1994, Abdel-Karim et al., 1997, Heikal et al., 1998).

iii) Calc-alkaline and tholeiitic compositions denote typical early intrusions during the transition from arc to continental magmatic activity and have generated from two magma sources (El-Metwally, 1992 and 1997, El-

Table 1. Results of the major, trace and REEs of some gabbros from south Sinai, Egypt

	Tholeiitic younger gabbros (THYG)														
	Wadi Rahaba				W. Tweiaba					El-Khamila					
	Hornblendite		Uralitized gabbro		Px-Hbl gabbro			Hbl gabbro		Oliv gabbro		Anorthosite		Gabbro	
	10 Rh2	11 Rh20	12 Rh7	13 Rh70	1-14 Tw5	2-15 Tw50	3-16 Tw51	4-17 Tw4	5-18 Tw40	19 SS8	20 SS80	21 SS3	22 SS30	23 SS3a	24 SS8a
SiO ₂	50.47	50.73	50.88	51.20	50.53	50.74	51.06	54.10	54.80	44.66	44.87	45.99	45.15	44.99	45.66
Al ₂ O ₃	15.05	14.85	14.63	14.58	16.62	16.43	16.58	15.12	15.10	15.35	15.54	17.13	16.50	16.13	15.35
Fe ₂ O ₃	12.46	12.08	12.39	11.87	12.59	12.94	12.48	11.62	11.18	17.08	16.12	15.14	15.77	16.14	17.58
MgO	7.11	6.58	6.75	7.12	5.11	5.09	5.14	4.57	4.35	6.70	6.02	7.03	7.31	7.03	5.70
CaO	9.85	10.11	10.16	9.94	8.29	8.31	8.24	7.11	7.13	8.44	8.43	9.19	9.32	9.19	8.44
Na ₂ O	2.13	1.97	1.94	1.98	2.95	2.94	2.88	2.60	2.54	2.54	2.51	2.21	2.17	2.21	2.54
K ₂ O	0.58	0.60	0.61	0.59	0.91	0.93	0.89	1.38	1.39	1.10	1.04	0.79	0.83	0.79	1.10
TiO ₂	1.08	1.03	1.04	1.06	1.49	1.45	1.51	1.58	1.52	1.78	1.69	1.49	1.43	1.40	1.58
P ₂ O ₅	0.23	0.24	0.24	0.25	0.28	0.29	0.27	0.37	0.39	0.64	0.59	0.17	0.18	0.17	0.64
MnO	0.25	0.23	0.23	0.25	0.22	0.23	0.24	0.19	0.18	0.27	0.31	0.15	0.15	0.15	0.27
LOI	0.45	0.98	0.68	0.27	1.00	0.51	0.64	1.60	0.99	1.05	0.95	0.65	0.63	0.65	0.55
Total	99.66	99.40	99.55	99.58	99.45	99.86	99.93	99.82	99.97	99.61	99.07	99.94	99.74	99.83	99.41
Cr	215	228	229	233	103	80	109	72	88	125	115	-	-	73	77
Ni	45	55	45	51	12	15	13	9	8	50	55	-	-	29	36
Ba	278	281	263	257	281	388	265	342	354	91	102	-	-	291	312
Th	3.1	2.8	2.2	1.2	3.0	2.9	3.2	3.8	4.2	2.6	2.1	-	-	2.9	4.0
U	1.2	1.8	1.1	2.4	0.5	1.1	0.8	0.4	0.3	1.1	2.8	-	-	3.6	4.2
Nb	7.4	7.5	7.3	7.9	5.6	4.3	4.4	9.1	10.5	3.7	3.8	-	-	6.9	7.5
Sr	450	525	486	501	383	312	397	295	302	696	681	-	-	802	745
Zr	71	90	81	74	73	62	77	64	69	70	65	-	-	62	71
Y	22	21	26	29	21	22	17	15	53	5	6	-	-	16	2
As	1.5	3	3	2	5.3	4.2	1.2	1.9	2.1	2.5	1.4	-	-	2	4
Co	52	34	41	44	30	33	20	28	27	37	34	-	-	45	30
Pb	11	8	12	15	8	8	4	10	10	5	5	-	-	9	6
Sc	55	23	47	51	47	49	53	46	48	23	25	-	-	20	23
W	2.7	3.9	3.3	4.1	0.0	0.6	0.3	0.8	0.7	1.8	1.5	-	-	1.3	2.1
Zn	48	112	95	102	94	99	74	87	90	38	42	-	-	114	18
La	-	18.60	20.15	-	15.50	-	-	17.05	-	9.30	-	-	-	16.12	-
Ce	-	24.30	40.50	-	20.25	-	-	24.30	-	13.67	-	-	-	29.16	-
Pr	-	3.60	6.00	-	2.40	-	-	3.25	-	1.98	-	-	-	3.84	-
Nd	-	19.20	30.00	-	12.00	-	-	15.60	-	7.20	-	-	-	12.00	-
Sm	-	4.00	6.40	-	3.80	-	-	5.20	-	1.30	-	-	-	1.60	-
Eu	-	1.48	2.22	-	1.90	-	-	2.00	-	0.56	-	-	-	0.59	-
Gd	-	5.72	3.12	-	2.60	-	-	3.12	-	1.09	-	-	-	1.17	-
Tb	-	0.55	0.65	-	0.54	-	-	0.65	-	0.12	-	-	-	0.17	-
Dy	-	3.52	3.84	-	2.56	-	-	3.52	-	0.48	-	-	-	0.83	-
Ho	-	0.77	0.84	-	0.49	-	-	0.63	-	0.11	-	-	-	0.15	-
Er	-	2.00	2.40	-	1.20	-	-	1.80	-	0.48	-	-	-	0.46	-
Tm	-	0.30	0.39	-	0.21	-	-	0.27	-	0.07	-	-	-	0.08	-
Yb	-	2.20	2.40	-	1.40	-	-	2.00	-	0.40	-	-	-	0.52	-
Lu	-	0.30	0.36	-	0.24	-	-	0.30	-	0.06	-	-	-	0.08	-
Σ	-	88.54	119.2	-	65.09	-	-	79.69	-	36.82	-	-	-	66.77	-

Table 1. Continue.....

Calc-alkaline younger gabbros (CAYG)									
	Imliq					W. Nakhil			
	Clino-pyroxenite		Gab.	Leucogabbro		Px-Hbl gabbro		Gabbro	
	14-1 Be27 a	15-2 Be22 b	16-3 Be25	17-4 Be27 b	18-5 Be22 a	6 Wt9	7 Wt90	8 Wt30	9 Wt3
SiO ₂	51.39	51.45	50.91	52.08	52.42	49.58	49.71	49.61	50.05
Al ₂ O ₃	19.31	16.45	17.86	18.51	16.40	18.41	18.44	18.32	18.28
Fe ₂ O ₃	7.71	8.24	7.96	7.84	8.20	7.91	7.69	8.11	8.21
MgO	6.37	7.92	7.13	7.05	7.89	6.69	6.59	6.71	6.45
CaO	8.17	10.24	9.28	8.85	9.39	10.00	9.97	10.02	9.84
Na ₂ O	3.73	2.78	3.30	2.98	2.87	2.38	2.34	2.32	2.48
K ₂ O	1.04	1.00	1.04	0.99	1.03	0.82	0.81	0.78	0.87
TiO ₂	0.85	0.49	0.66	0.67	0.47	0.53	0.54	0.55	0.58
P ₂ O ₅	0.04	0.05	0.09	0.09	0.03	0.12	0.11	0.12	0.10
MnO	0.14	0.16	0.15	0.15	0.15	0.12	0.13	0.11	0.13
LOI	0.95	1.02	1.17	0.55	0.70	3.27	3.24	3.02	2.44
Total	99.70	99.80	99.52	99.76	99.55	99.83	99.57	99.67	99.43
Cr	110	155	108	111	149	83	78	141	88
Ni	46	55	44	39	52	44	58	39	69
Ba	406	199	425	242	1 84	269	275	250	195
Th	1.8	0.8	1.6	2.3	0.7	2.6	3.1	2.3	2.8
U	4	2.8	4.2	4.4	3.2	2.6	2.1	0.9	1.6
Nb	4.9	3.7	5.1	4.5	3.6	5.5	7.1	5.1	5.9
Sr	746	644	734	781	633	724	781	694	712
Zr	42	11	57	43	30	77	81	77	77
Y	14	7	15	13	7	12	14	10	13
As	0	0	0	0	0	0.7	0.9	0.3	0
Co	28	44	30	25	39	43	55	46	23
Pb	9	6	7	8	7	11	18	2	10
Sc	28	41	29	26	42	24	19	21	20
W	2.9	2.7	3.1	1.5	2.5	1.8	2.1	1.1	1.6
Zn	71	135	71	111	211	71	66	72	46
La	8.06	-	8.40	-	5.60	16.43	-	-	13.95
Ce	26.73	-	23.50	-	13.77	26.73	-	-	21.06
Pr	3.24	-	2.40	-	2.04	3.72	-	-	2.52
Nd	16.80	-	14.40	-	7.80	16.80	-	-	15.00
Sm	2.80	-	2.60	-	1.80	2.40	-	-	4.20
Eu	1.12	-	1.04	-	0.59	0.84	-	-	1.12
Gd	2.47	-	1.69	-	1.56	2.34	-	-	3.90
Tb	0.43	-	0.30	-	0.26	0.35	-	-	0.06
Dy	2.30	-	1.66	-	1.38	1.82	-	-	2.88
Ho	0.43	-	0.35	-	0.29	0.33	-	-	0.60
Er	1.26	-	1.20	-	0.94	1.00	-	-	1.60
Tm	0.20	-	0.18	-	0.16	0.14	-	-	0.23
Yb	1.24	-	1.08	-	0.90	1.02	-	-	1.56
Lu	0.17	-	0.16	-	0.10	0.14	-	-	0.22
Σ	67.25	-	58.96	-	37.29	74.06	-	-	68.90

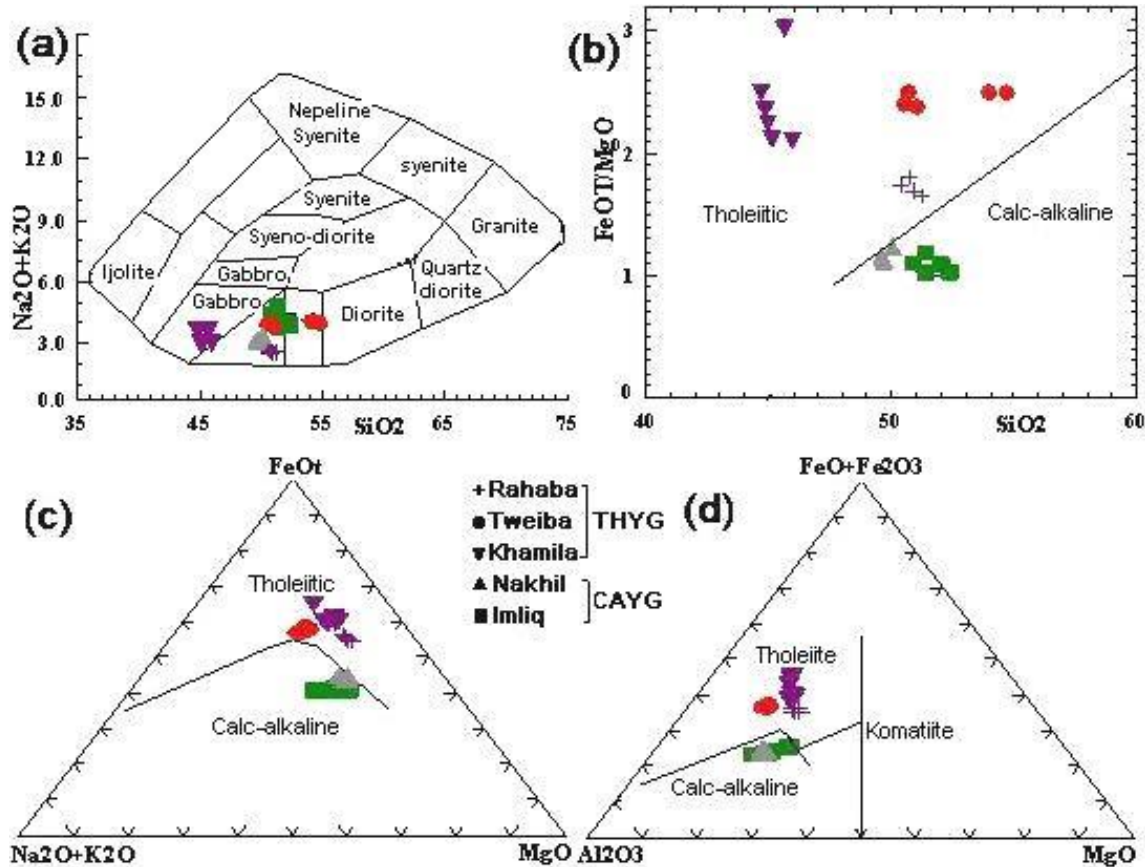


Figure 3. Variation diagrams showing the chemical classification of the studied gabbros, a) b) Alkalis versus silica diagram adapted by Wilson (1989); c) FeOT/MgO versus SiO₂ of Miyashiro (1974); d) AFM diagram of Irvine and Baragar (1971); e) FeO+Fe₂O₃-Al₂O₃-MgO diagram of Jensen (1976).

Sheshtawi et al., 1994).

On the other hand, El-Mahallawi (1996) mentioned that gabbroic rocks of El-Ginena El Gharbia exhibit a definite calc-alkaline nature and have a transitional character from uncontaminated within plate basalts to crustal contaminated mafic rocks of active continental margin indicating their derivation from a mantle source which experienced subduction.

The analyzed samples of the present younger gabbros display two different trends with tholeiitic and calc-alkaline affinities and their averages chemical composition (Table 1) are analogous to those mentioned by El-Metwally (1997) and Moghazi (1994). The calc-alkaline younger gabbros (CAYG) are generally higher in SiO₂, CaO, Ni and Sr, and lower in FeO[†], TiO₂, Al₂O₃, MnO, Zr, Y, Sc and total REE content compared to the tholeiitic younger gabbros (THYG). It is notable that the THYG are exposed at Wadi Rahaba, Wadi Tweiba and El-Khamila. While the CAYG are represented by those exposed at Imliq and Wadi Nakhil. On SiO₂ versus alkalis diagram, all the analyzed gabbroic rocks plot mainly in the gabbro field with minor gabbro-diorite field (Figure 3a).

Geochemical Classification and Magma Type:

The analyzed samples reveal again two different trends for the present gabbroic rocks tholeiitic and calc-alkaline. The analyzed samples of Imliq and W. Nakhil areas (CAYG) display calc-alkaline nature, while those of Wadi Rahaba, Wadi Tweiba and El-Khamila areas (THYG) show tholeiitic tendency (Figure 3b,c,d). The same feature was concluded for the younger gabbros in Sinai (El-Metwally, 1986, 1992 and 1997, Moghazi, 1994, Heikal et al., 1998).

Good correlation of the major and trace elements against FeOT/MgO ratio are also evident on the Harker diagrams (Figure 4) which used to differentiate between the THYG and CAYG as well as to explain the mobility and some genetic relations. This figure reveals a compositional gap amongst the gabbroic rocks related to their tholeiitic and calc-alkaline tendencies.

The THYG are characterized by increasing FeOT/MgO ratio with increasing TiO₂, MnO, Na₂O, K₂O, P₂O₅, Y, Th and As and decreasing Al₂O₃, CaO, MgO, Cr, Ni, Co, Zn contents. The noticeable scatter (especially Zr, Y) can be explained by varying amounts of crystal fractionation and

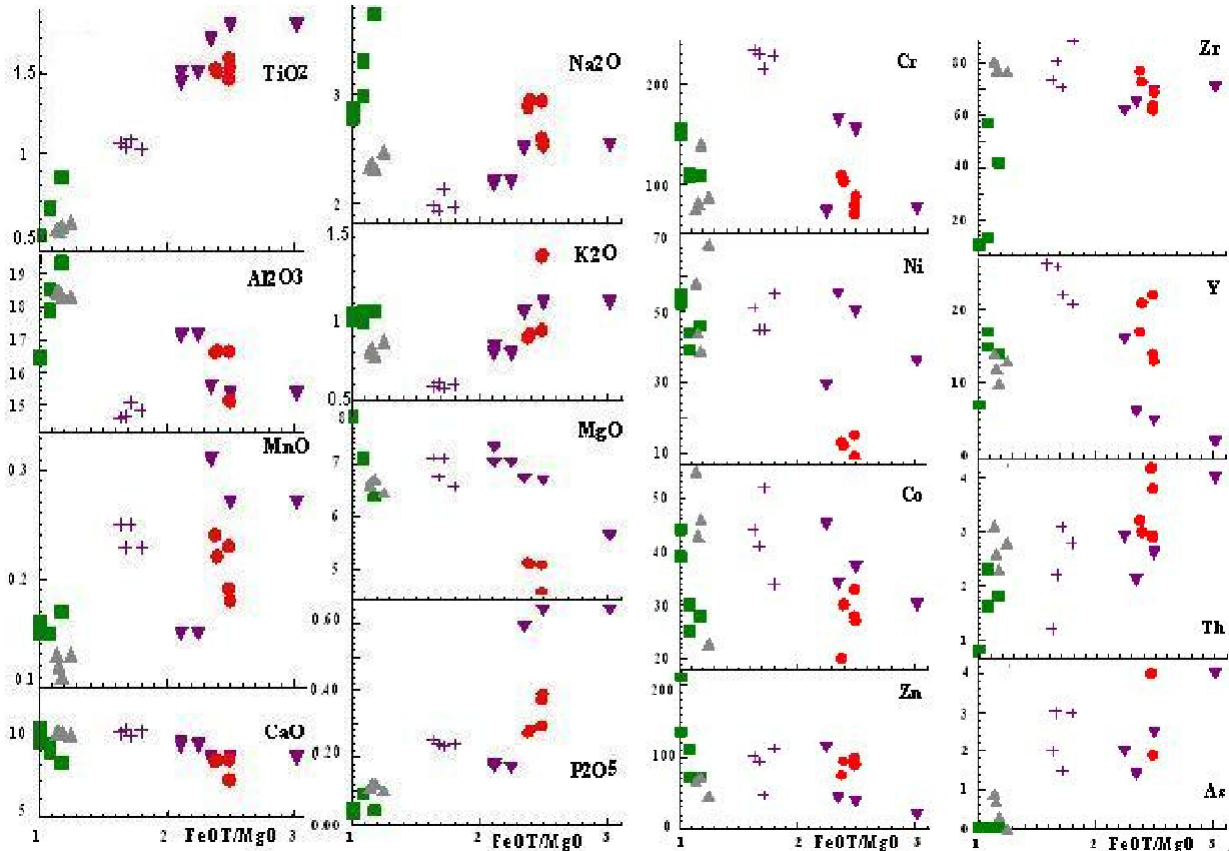


Figure 4. Harker variation diagrams for major (Wt.%) and trace elements (ppm) versus FeOT/MgO.

differentiation. Apart the scattered distribution, the contents of TiO_2 , Zr and Th increase, while those of Al_2O_3 , CaO, MgO, Cr and Zn decrease with increasing FeOT/MgO ratio in the CAYG. These features indicate the role of fractionation of clinopyroxene, plagioclase, olivine and Fe-Ti oxides for present gabbros. Sr behaves incompatibility with differentiation suggesting plagioclase accumulation for the tholeiitic-suites but unlikely in calc-alkaline ones (DeBari, 1994).

Crustal Contamination and Tectonic Setting

The role of crustal contamination must be considered in all intraplate magmatic environments (Wilson, 1989). Some elemental ratios (e.g., Y/Nb, Nb/Zr and La/Zr) are utilized as parameters for crustal contamination. Weaver et al. (1972) and Lippard (1973) reported that if these ratios are low or constant throughout the rocks, then the crustal contamination is unlikely to have occurred.

In the CAYG of W. Nakhil and Imliq, the above mentioned ratios are nearly constant or of narrow range values (4.5-10.0 Y/Nb, 0.02-0.05 Nb/Zr and 0.15-0.22 La/Zr). A feature indicates that crustal contamination played a minor role throughout these gabbros. On the

other hand, the elemental ratios Y/Nb (2.1-16.9), Nb/Zr (0.01-0.11) and La/Zr (0.21-0.55) are variable throughout the THYG of W. Tweiba, W. Rahaba and El-Khamila which indicate dominant crustal contamination upon these rocks. Moreover, the existence of more prominent negative Nb-anomalies of the CAYG with respect to that of the THYG may have been reflected a more crustal contamination of the former gabbros, because the LIL elements are typically enriched compared with Nb in the continental crust (Taylor and McLennan, 1985).

The plots of the investigated gabbros fall in the continental field on the Zr-Zr/Y diagram of Pearce (1983) (Figure 5a). On the $\text{FeO}^t\text{-MgO-Al}_2\text{O}_3$ diagrams of Pearce et al. (1977) (Figure 5b), the CAYG lie within the island arc (or volcanic arc) and active continental margin field, whereas the THYG fall in the continental field. This feature suggests a transitional regime towards the intraplate tectonic setting. The same conclusion was concluded for the younger gabbros in Sinai (El-Metwally, 1997). However, samples lie in the orogenic field indicating that these rocks were generated originally by varying degree of crystal fractionation and differentiation of parental magma derived from upper mantle –lower crust materials (Jubeli and Heikal, 1994, Heikal et al., 1998).

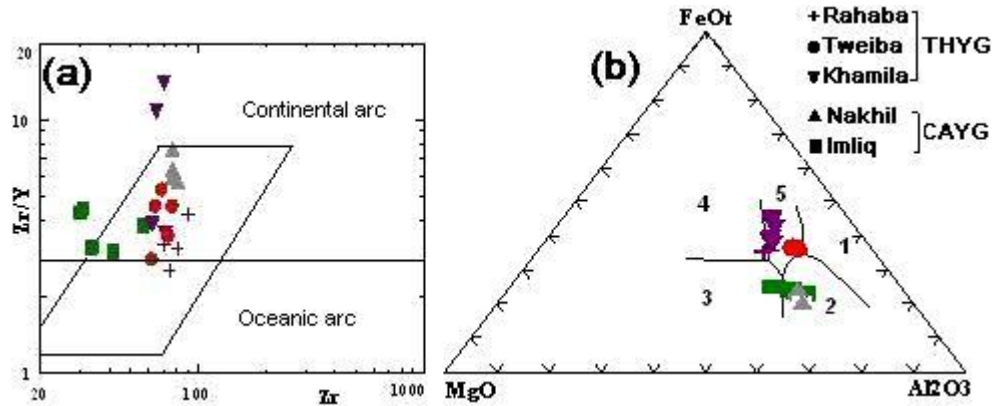


Figure 5. a) Zr/Y vs. Zr diagram of Pearce (1983) for the studied gabbros; b) FeO^{\dagger} -MgO- Al_2O_3 diagram of Pearce et al. (1977). 1: Spreading center island, 2: orogenic (island arc and active continental margin), 3: Oceanic ridge and floor, 4: Oceanic island, 5: continent.

REEs

The chondrite normalized REE patterns (Table 1, Figure 6) show that present gabbros are enriched in REE relative to the chondritic abundances, and enriched in LREE (La-Sm) relative to HREE (Gd-Yb). Generally, they have moderately fractionated REE patterns. The patterns indicate that there are differences between the tholeiitic (THYG) and calc-alkaline younger gabbros (CAYG). The THYG have a wider range of REE with respect to that of the CAYG. The overall abundance of REE varies from 1 to 70 times chondritic in the THYG compared with that of the CAYG (3-50 times) (Figure 6a,b).

The THYG have moderately fractionated REE pattern where $(\text{Ce}/\text{Yb})_n$ varies from 2.7 to 13.8 and $(\text{La}/\text{Lu})_n$ varies from 5.4 to 20.8. The REE patterns of both gabbros from W. Rahaba and Tweiba are rather similar. El-Khamila gabbros are characterized by more depleted moderate REE [$(\text{Sm}/\text{Yb})_n = 3.1$] and HREE [$(\text{Gd}/\text{Lu})_n = 1.8-2.0$] abundance with respect to that from both W. Rahaba and W. Tweiba [$(\text{Sm}/\text{Yb})_n = 1.8-2.7$] and HREE [$(\text{Gd}/\text{Lu})_n = 1.0-2.2$] (Figure 6a) probably indicate their less fractionalization. Moreover, a progressive increase in REE abundance of the present THYG (as well as their fractionation trend) from El-Khamila through W. Tweiba to W. Rahaba is recorded.

The CAYG have moderately fractionated REE pattern where $(\text{Ce}/\text{Yb})_n$ varies from 3.3 to 6.5 and $(\text{La}/\text{Lu})_n$ varies from 4.7 to 11.0 (Figure 6b). The REE patterns of both gabbros from W. Nakhil and Imliq are rather similar. However, a progressive increase in REE abundance of the present CAYG (as well as their fractionation trend) from Imliq to W. Nakhil is recorded.

A slight positive Eu anomaly is observed in some samples from the studied gabbros of all areas due to plagioclase accumulation. The normalized REE patterns of the average values of all studied gabbros (Figure 6c) exhibits a progressive evolution from El-Khamila through

Imliq, W. Nakhil, W. Tweiba to W. Rahaba suggesting their fractionation trend.

The averages of studied gabbros are plotted on MORB-normalized patterns (Figure 6d) compared with basalts from different sources (Sun, 1980). On this diagram, the arc related environment is effectively distinguishable from the within plate setting on the basis of a diagnostic Nb, Ce and Ti trough (Pearce, 1983, Tepper, 1996, etc.) reflect the so-called a subduction zone component. The diagram also has the advantage of illustrating -at a glance- the degree of arc maturity. From figure (6d) it is interesting to note that the present suites occupy the intermediate portion between the arc and continental margin basalts and seem to represent a transitional stage between a continental island arc and a thick continental margin as well as reflect a thin continental margin. However, the CAYG are characterized by higher U and Sr and lower Nb, Y and Tb contents with respect to the THYG probably due to their less affect by subduction and transition to continental margin setting.

DISCUSSION AND PETROGENETIC ASPECTS

In terms of petrogenesis it is interesting to speculate on the process (es) which explain the observed variation throughout the suits.

In spite of the investigated THYG and CAYG are closely related in space and time, they do not form a single comagmatic suite and may have been intruded over a wide period of time. Similar conclusion was given for the younger gabbros in Sinai (El-Metwally, 1997). Nevertheless, diversities in chemistry between both varieties are mentioned above.

The ratios between incompatible elements (e.g., Zr/Nb, Zr/Y and Zr/Sr) vary significantly amongst the present gabbros to extend that they can not theoretically be explained by fractional crystallization alone. Therefore,

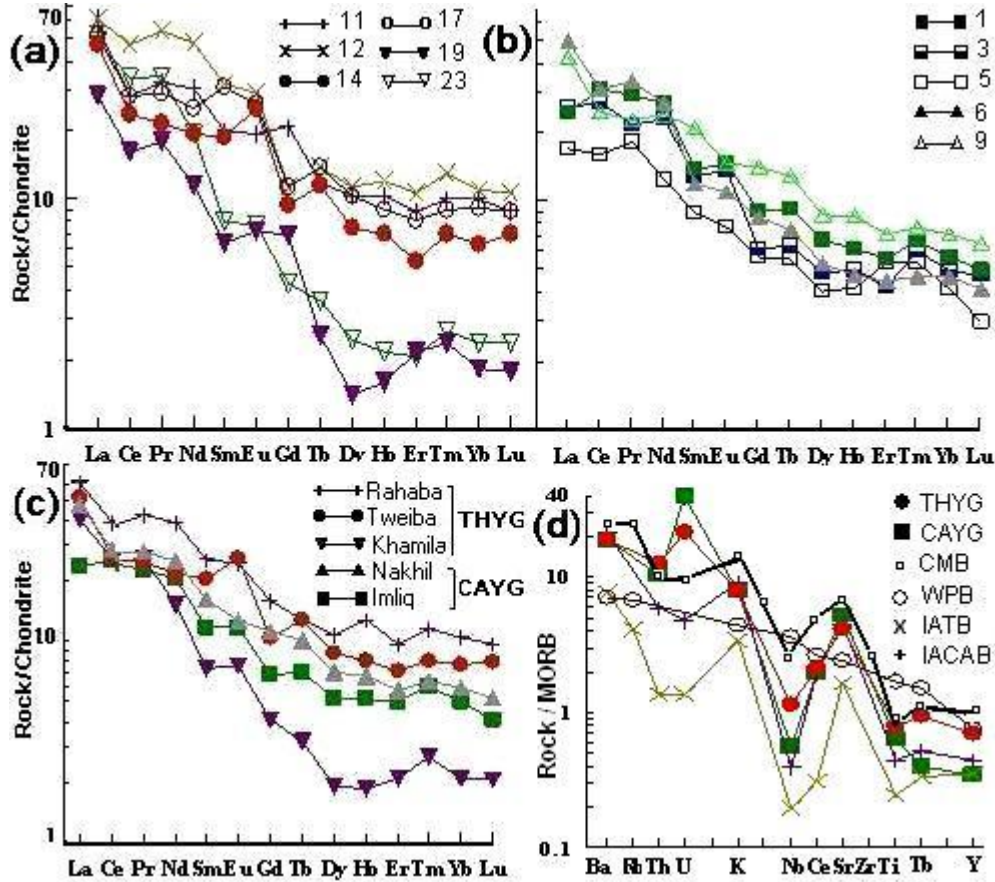


Figure 6. Chondrite-normalized REE patterns for: a) Tholeiitic younger gabbros (THYG); b) Calc-alkaline younger gabbros (CAYG); c) Averages of younger gabbros varieties. d) MORB-normalized patterns of the studied gabbros compared with continental margin (CMB), within plate (WPB), island arc tholeiite (IATB) and island arc calc-alkaline basalts (IACAB) of Sun (1980). The normalizing values after Pearce (1982).

the expected calc-alkaline magma chamber was open to periodic enrichment by primitive melt, since the clinopyroxenites cut the gabbros, as well as the existence of rhythmic layering in gabbros. Differentiation was controlled by cotectic crystallization of ortho-, clinopyroxene and plagioclase followed by the interface of the pyroxene with residual liquid to form amphibole. Moreover, the presence of the hornblendites and hornblende-rich layered and normal gabbros suggest that the hydrous phases are crystallized according to the increase in volatile activity due to the early crystallization of anhydrous phases (olivine and pyroxene) as it has been proved experimentally by Baker and Egger (1983). The hydrous phases are formed by reaction between the crystallized anhydrous phases and the liquid. Similar result was obtained by El-Metwally (1997) for the younger gabbros in Sinai.

The low FeOT/MgO (<3.1), and the low values of Cr (<220 ppm), Ni (<70 ppm) and Co (<54ppm) of the present gabbros indicate that the considered magma was not primary and was not directly formed by partial melting, but has experienced varying degrees of crystal

fractionation and differentiation of olivine and pyroxene. Partitioning of Cr and Ni into the fractionated olivine and clinopyroxene are evident in Figure 7. The high FeOT/MgO ratios of the tholeiitic samples indicate that olivine and clinopyroxene fractionation have occurred before the emplacement. In addition, crustal contamination plays an important role in modifying the chemical affinities of the THYG from El-Khamila, W. Tweiwa and W. Rahaba.

The THYG are characterized by high contents of Fe₂O₃T, TiO₂, Zr and FeOT/MgO ratios. The enrichment FeO[†] and TiO₂ in the present gabbros is explained by: i- a low degree partial melting (Dupuy et al., 1988); ii- crustal contamination (Heikal et al., 1988) and/or iii-subduction related contamination by subcontinental lithosphere (Murphy, 1988). The crustal contamination is considered as an important parameter in the formation of continental tholeiites (Cox and Hawkesworth, 1984).

Significant differences between the THYG and the CAYG exist in Figures (3, 4, 5). These features reflect the dissimilarities in the parental magmas composition for each of them. In general, the CAYG are relatively higher

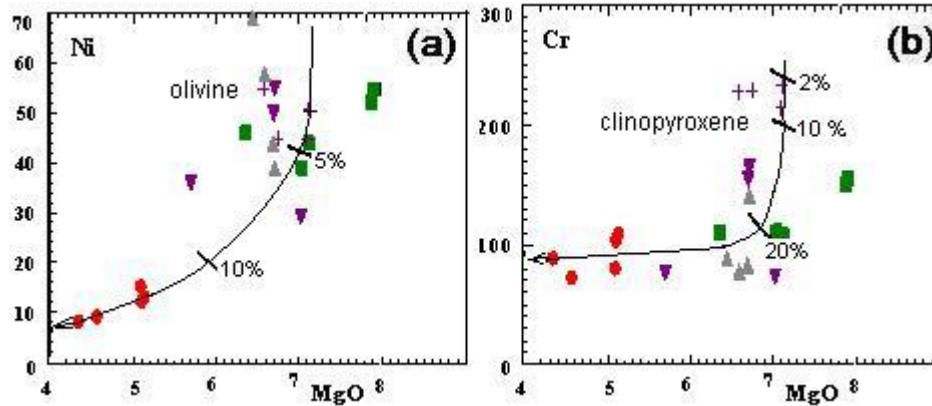


Figure 7. Variation in Ni and Cr contents in relation to MgO concentration. Trends and degrees of fractionation of olivine and clinopyroxene are shown for comparison.

in SiO₂, CaO, Ni and Sr, and lower in FeO[†], TiO₂, MnO, Zr, Y, Sc, total REE contents and FeO[†]/MgO ratios compared to the THYG. The melt of the CAYG was probably accompanied by high oxygen fugacity and water activity. The water rich melt crystallized a hydrous mineral such as hornblende and biotite instead of pyroxene. The present gabbros were suffered a minor crustal materials involvement.

Based on the REE distribution, the gabbroic rocks can be grouped into 3 types (Zimmer et al., 1995), namely i-Plag-rich gabbros have LREE enriched and HREE depleted patterns with distinct positive Eu anomaly, ii-Plag-cpx gabbros have a relatively flat REE pattern with no Eu anomaly and iii-Gabbros with less plag and a higher proportion of cpx have LREE depleted patterns with no Eu anomaly.

Most investigated gabbros have LREE enrichment, HREE depletion and a small positive Eu anomaly (Figure 6). These patterns result from dominant proportion of plag with respect to the cpx in the samples as also indicated by petrographic observations. Moreover, the absence of a positive Eu anomaly in some samples which have cumulus plagioclase is interest. This feature is probably due to either the variation in oxidation state of Eu during fractionation which may control the extent to which Eu partitions into plagioclase or attributed to pressure fractionation before emplacement (Henderson, 1984).

The present gabbros seem to be generated by varying degree of crystal fractionation and differentiation of parental magma derived from heterogeneous (upper mantle-lower crust) materials. They occupy the intermediate portion between the arc and continental margin basalts and seem to represent a transitional stage between a continental island arc and a thick continental margin. However, the CAYG are characterized by higher U and Sr and lower Nb, Y and Tb contents with respect to the THYG probably due to their less affect by subduction and by transition to continental margin setting.

CONCLUSION

The present work focuses on five minor gabbroic intrusions occurring in three main areas: Southeastern Sinai (W. Tweiba, 3.2 km² and W. Nakhil, 1.0 km²), Central Sinai (W. Rahaba, 4.2 km²) and Southeastern Sinai (Imliq, 40 km² and El-Khamila, 13 km²). Their contacts with the enveloped granitoid rocks are mostly transitional to knife-sharp of intrusive nature. Absence of chilled margins along their contacts is ascribed to emplacement from shallow depths. Rhythmic layers are recorded in Imliq and El-Khamila.

The present gabbros are mainly represented by normal gabbro, pyroxene-hornblende gabbro and hornblende gabbro with uncommon clinopyroxenite, hornblendite, olivine gabbro, uralitized gabbro, leucogabbro and anorthosite. The corresponding textures of these rocks are dominantly mesocumulate with scarce orthocumulate, and adcumulate type (Irvine et al., 1998).

The studied gabbroic rocks were derived from tholeiitic (El-Khamila, W. Tweiba and W. Rahaba occurrences) and calc-alkaline (Imliq, W. Nakhil occurrences) magma (Figure 3, 4, 5). Both two magma types are closely related in time and space, but with different geochemical trends as well as the tectonomagmatic setting. The low FeO[†]/MgO ratios and the low values of Cr, Ni and Co of both types indicate that they were not primary magma and were not directly formed by partial melting, but suffered from fractionation before emplacement.

The tholeiitic younger gabbros (THYG) of W. Tweiba, W. Rahaba and El-Khamila areas are generally characterized by high FeO[†], TiO₂, MnO, Zr, Y, Nb, Th and total REE contents as compared calc-alkaline younger gabbros. The association of the hornblendites and hornblende-rich layered with normal gabbros in THYG suggest that the hydrous phases are formed by reaction between the crystallized anhydrous phases (olivine and pyroxene) and volatiles due to an early crystallization. The parental magma of the THYG was controlled by

multistage processes including varying amounts of crystal fractionation and differentiation with prominent crustal contamination.

The calc-alkaline younger gabbros (CAYG) of Imliq and W. Nakhil areas are characterized by high contents of SiO₂, Al₂O₃, CaO, Sr and U with respect to the THYG. The significant variation of incompatible elements ratios (e.g., Zr/Nb, Zr/Y and Zr/Sr) amongst the present CAYG indicates that the fractional crystallization was not the only process of formation. The expected calc-alkaline magma chamber was periodically enrichment by primitive melt, since the clinopyroxenites cut the gabbros, as well as the existence of rhythmic layering in gabbros. The parental magma of the CAYG was controlled by Differentiation and crystallization of pyroxene and plagioclase which accompanied by high oxygen fugacity and water activity. The water rich magma was crystallized hornblende and biotite instead of pyroxene. The present gabbros were suffered a minor crustal materials involvement.

The existence of the present two magma types (tholeiitic and calc-alkaline) reveals a heterogeneous mantle source. The diversities of both kinds of gabbros are attributed to the varying degrees of partial melting, crustal contamination and/or subduction related contamination by subcontinental lithosphere.

The investigated gabbros of all localities have LREE enrichment; HREE depletion and a small positive Eu anomaly suggest their moderate differentiation. The REE patterns reveal the dominant role of plagioclase over the clinopyroxene as also indicated by petrographic observations. Moreover, the absence of a positive Eu anomaly in some samples which have cumulus plagioclase is due to pressure fractionation before emplacement. A slight positive Eu anomaly is observed in some studied gabbros attributed to plagioclase accumulation. The normalized REE patterns reveals a progressive fractionation from Imliq to W. Nakhil gabbros in CAYG and from El-Khamila through, W. Tweiwa to W. Rahaba gabbros in THYG.

The studied gabbros were mostly generated and emplaced in continental crust and tend to be formed by a transitional regime passing from the final stage of arc compression to extensional active continental margin or seem to represent a transitional stage between a continental island arc and a thick continental margin.

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